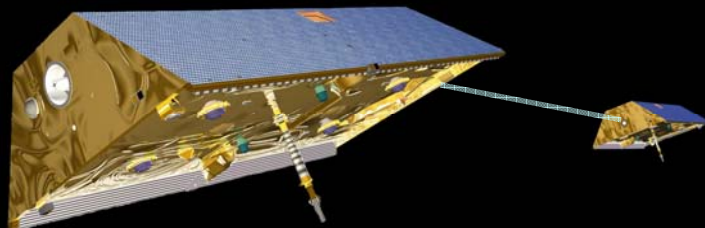


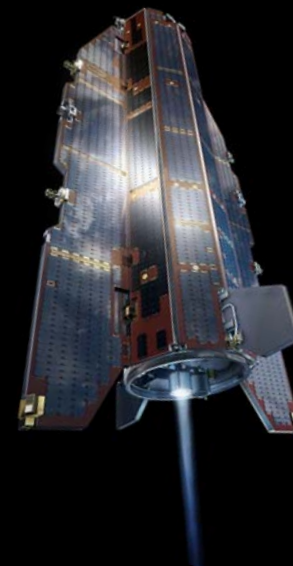
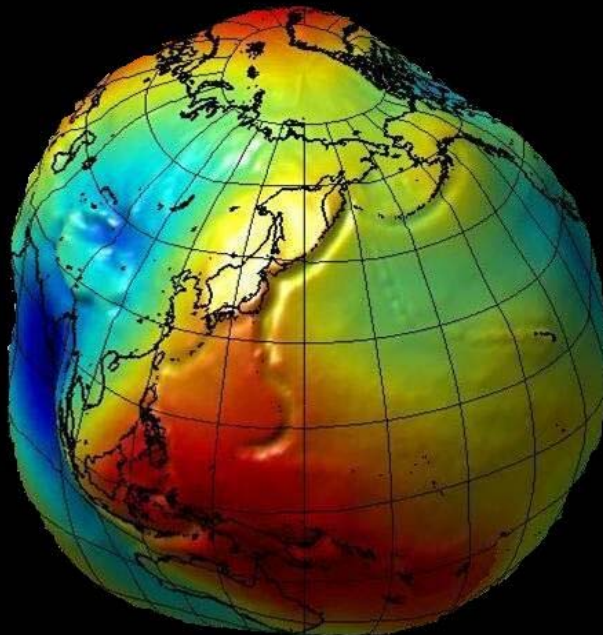
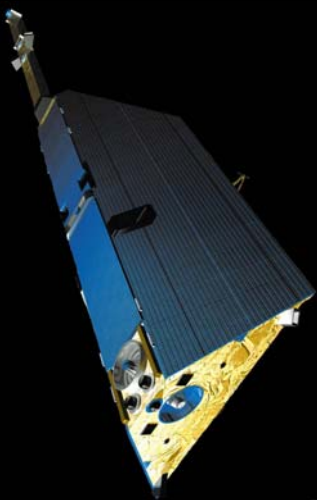
# Earth Gravity Field Determination from Space – A Computational Challenge

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Adrian Jäggi<sup>2</sup>, Reiner Rummel<sup>1,2</sup>, Lieselotte Zenner<sup>1</sup>

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Physikalische Geodäsie,  
Technische Universität München.

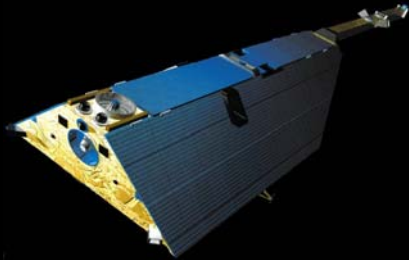


<sup>2</sup>Institute for Advanced Study,  
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& Astronomisches Institut,  
Universität Bern.

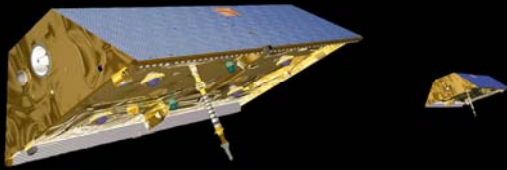


# Outline

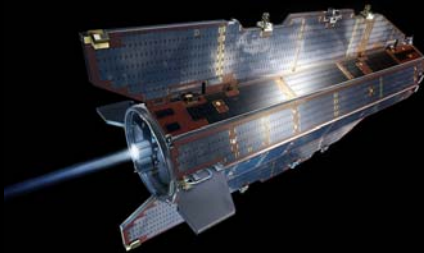
1. The Earth Gravity Field:  
A few Basics
2. How do we Observe  
Gravity from Space?
3. Computational Challenges
4. Summary & Conclusions



CHAMP



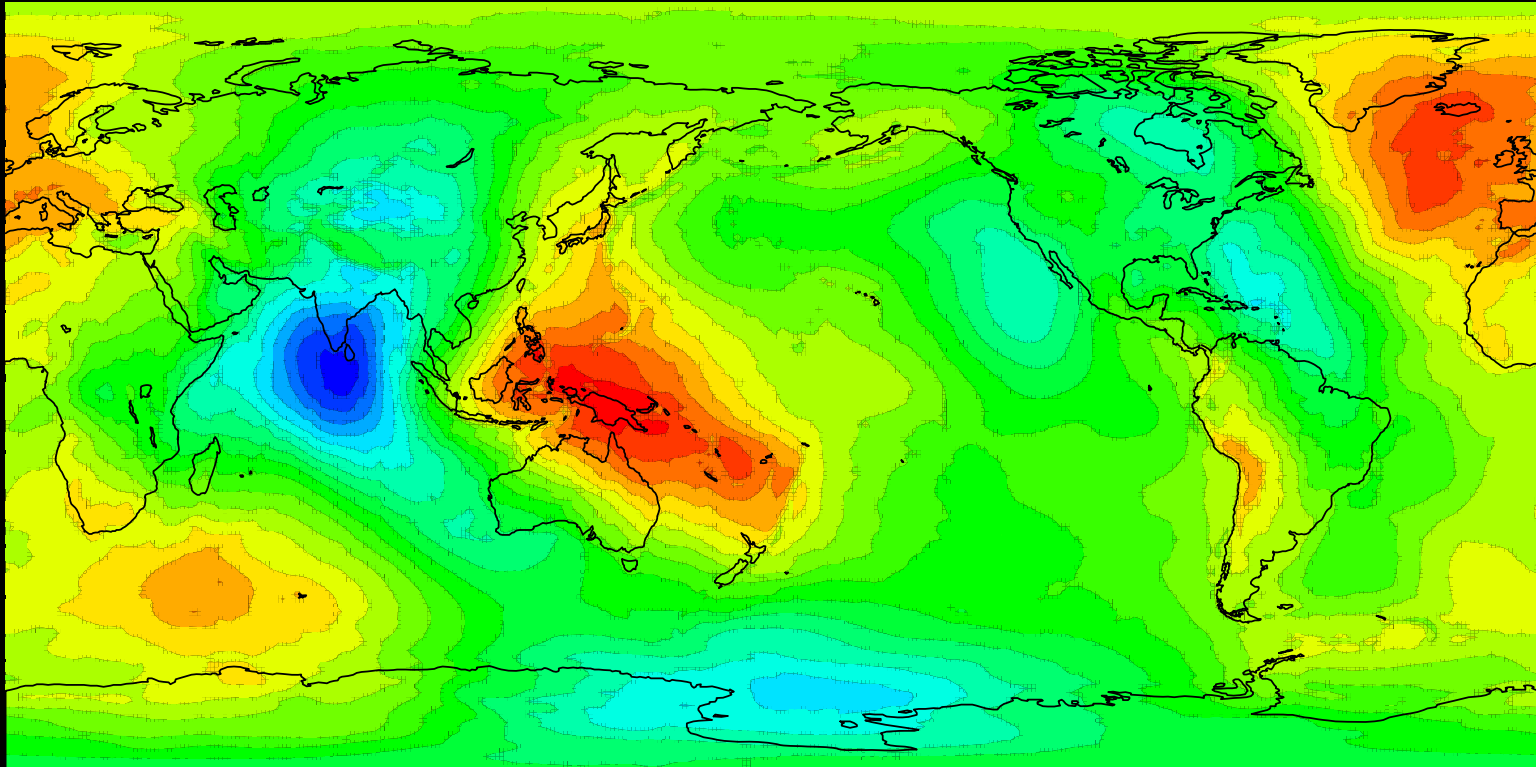
GRACE



GOCE

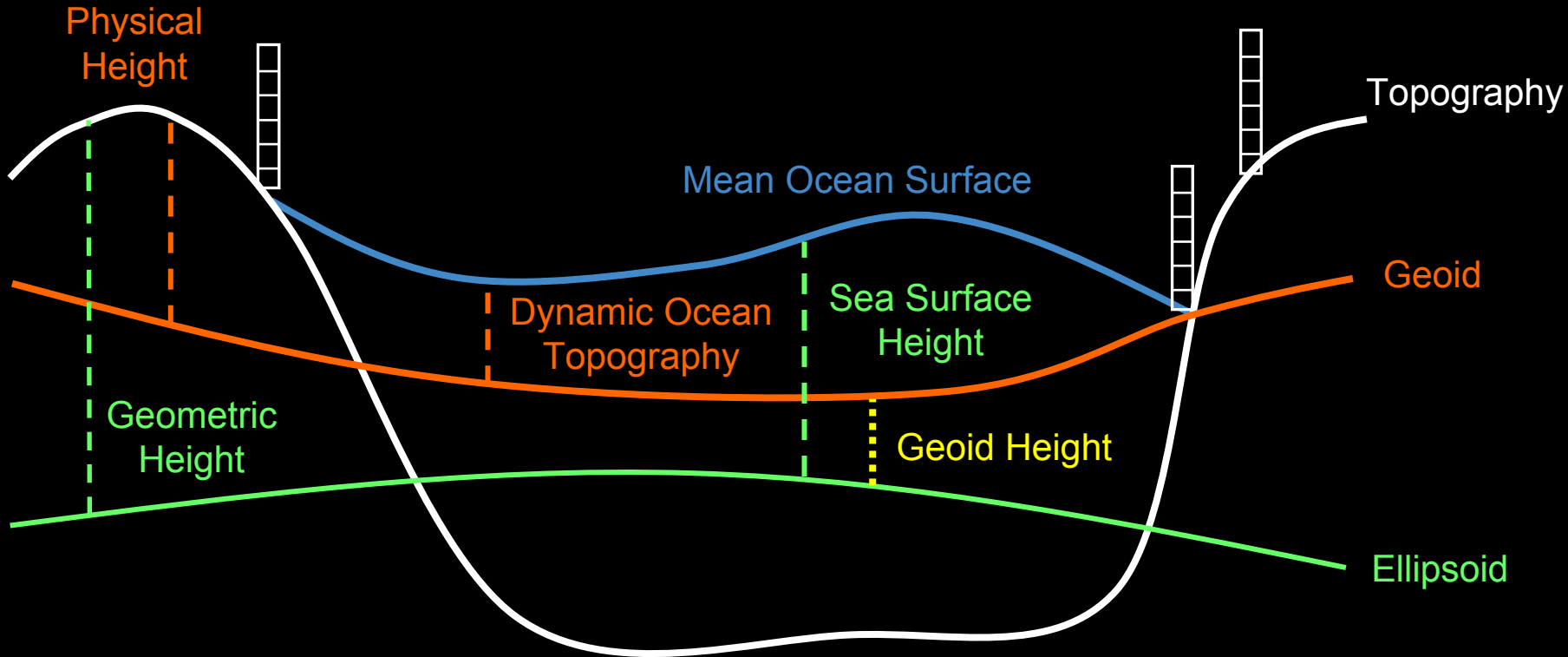
# The Earth Gravity Field

## The Geoid



# The Earth Gravity Field

## The Geoid



**Equipotential Surface = Surface of equal Gravity**

**Reference of Physical Height Systems**

# The Earth Gravity Field

## Mathematical Model

Disturbing  
Potential

$$T = \frac{GM}{r} \sum_{n=0}^N \left(\frac{a}{r}\right)^n \sum_{m=0}^n (C_{nm} \cos m\lambda + S_{nm} \sin m\lambda) P_{nm}(\cos \theta)$$

Geoid  
Height

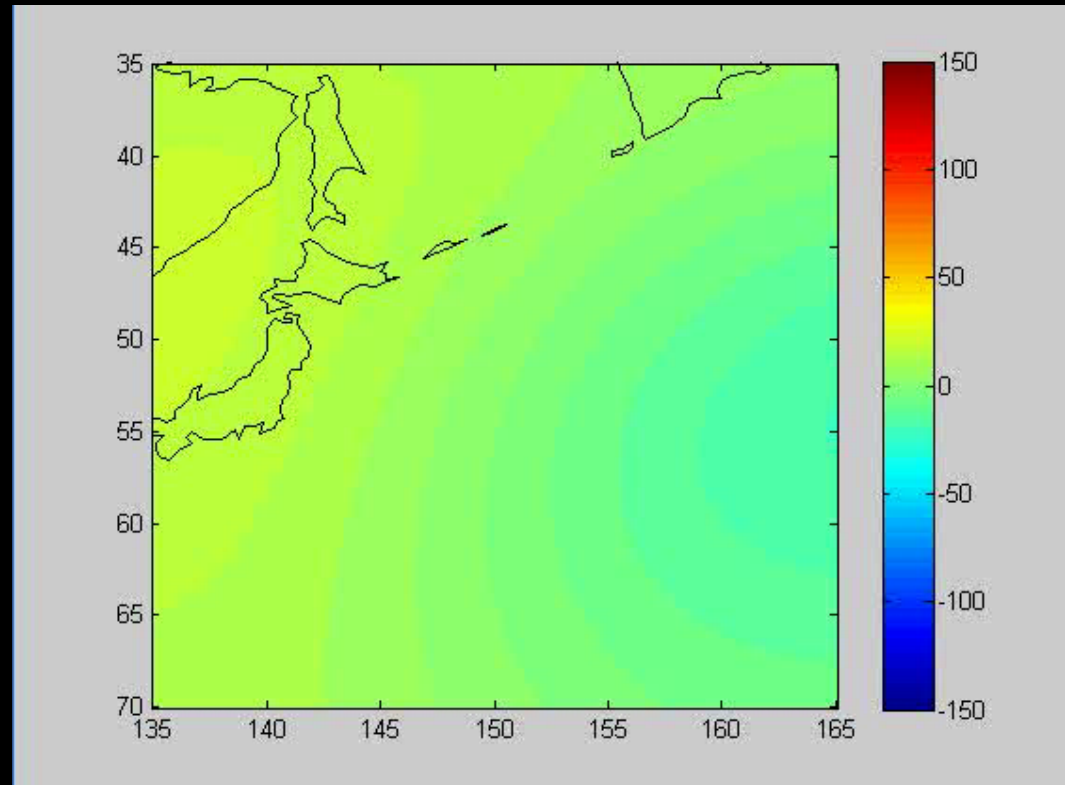
$$N = \frac{T}{\gamma}$$

Gravity  
Anomaly

$$\Delta g = -\frac{\partial T}{\partial r} - \frac{2}{r}T$$

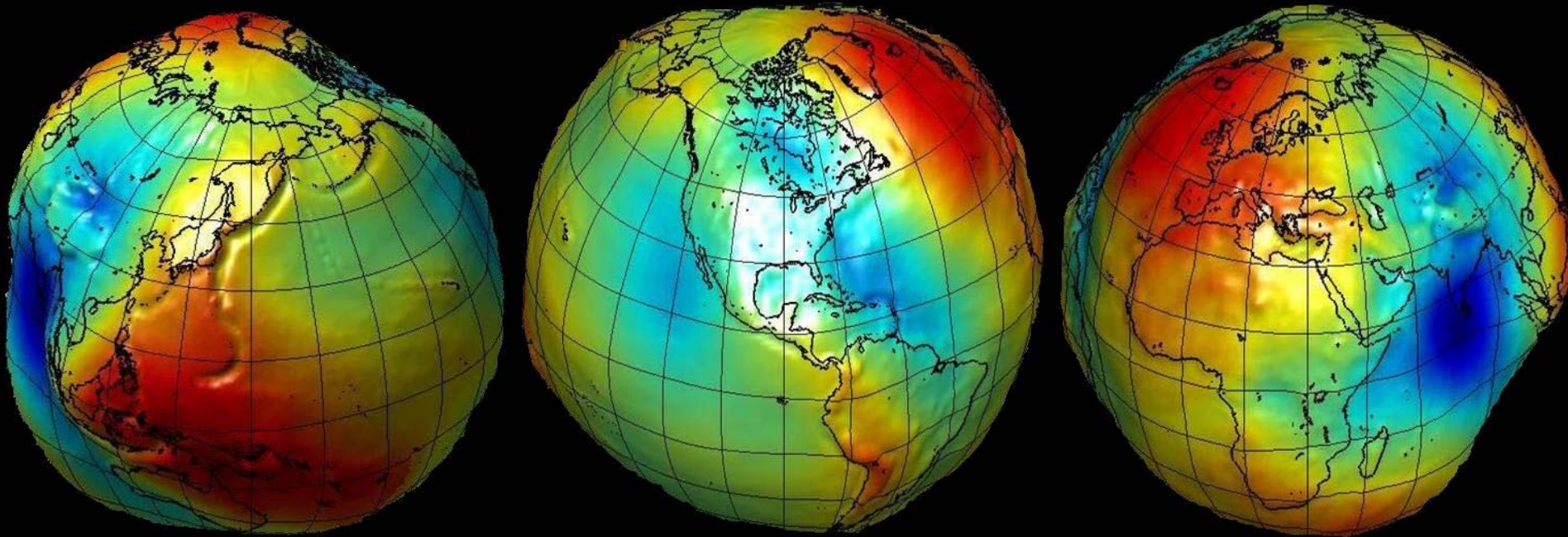
Gravity  
Gradient

$$T_{rr} = \frac{\partial^2 T}{\partial^2 r}$$



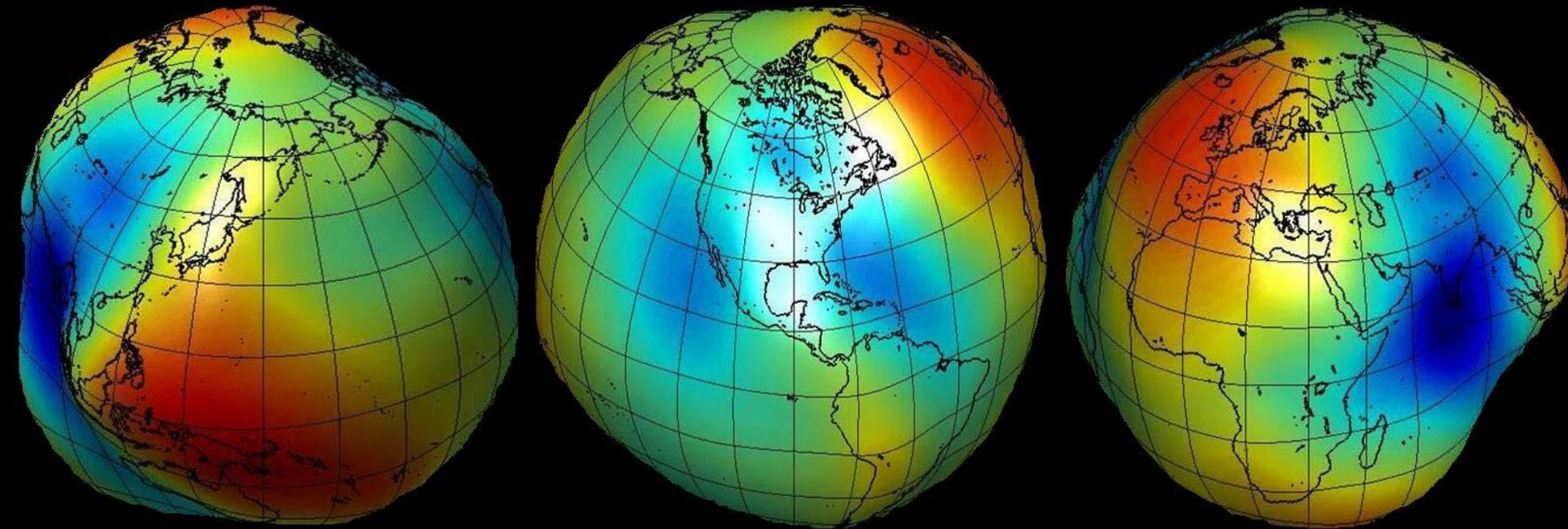
# The Earth Gravity Field

Seen on Ground



# The Earth Gravity Field

Seen in 500 km Height



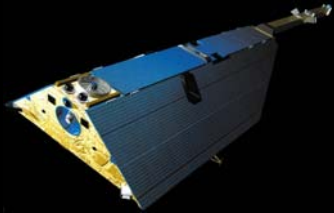
# Observing Gravity Field from Space

## Boundary Conditions of Mission Design

1. A satellite can be regarded as a free-falling test mass in the Earth gravity field.
2. Due to quadratic attenuation of gravity signal with height the satellite has to fly as low as possible.
3. Low orbits imply higher impact of non-gravitational forces, which have to be modeled or observed and/or compensated.
4. The orbit of a satellite (positions and velocities) represents an indirect gravity field observation. Therefore continuous positioning of satellite is required.
5. Signal attenuation partly can be compensated by observing derivatives of the gravity potential in space.

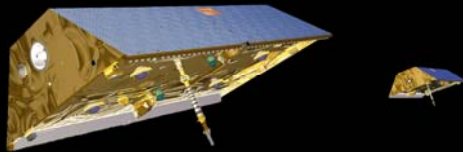
# Observing Gravity Field from Space

## Status of Satellite Missions



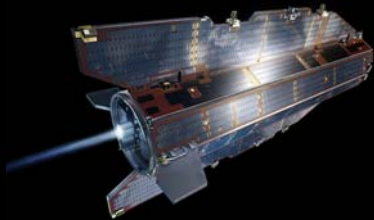
**CHAMP**

**Launch:** 15.7.2000  
**Funded by:** GFZ Potsdam / DLR  
**Lifetime:** 8 years (extension possible)  
**Orbit:** 87 degree inclination, 450-300 km height  
**Goals:** static gravity & magnetic field, atm. sounding



**GRACE**

**Launch:** 17.3.2002  
**Funded by:** NASA / DLR, Science CSR Austin & GFZ  
**Lifetime:** 8 years (extension possible)  
**Orbit:** 89 degree inclination, 450-300 km height  
**Goals:** static & time variable gravity field



**GOCE**

**Launch:** planned for 1. quarter 2009  
**Funded by:** ESA  
**Lifetime:** 2 years (extension possible)  
**Orbit:** 96.5 degree inclination, 260-270 km height  
**Goals:** static gravity field

# Observing Gravity Field from Space

## Satellite-to-Satellite Tracking – High Low Case

Observation of the low Frequencies

GPS Constellation

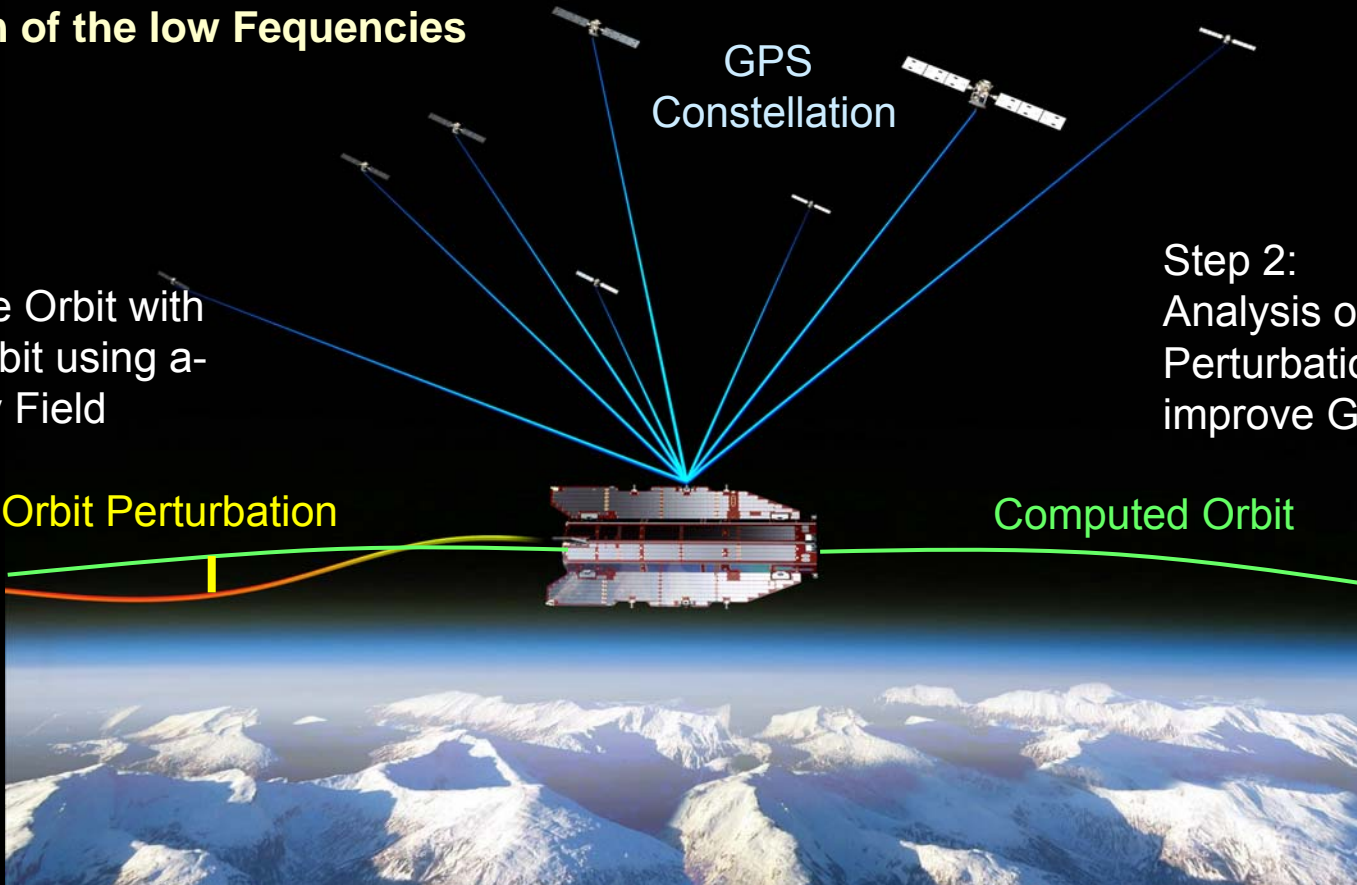
Step 1:  
Compare true Orbit with  
computed Orbit using a-  
priori Gravity Field

Step 2:  
Analysis of Orbit  
Perturbations to  
improve Gravity Field

Orbit Perturbation

Computed Orbit

True Orbit



**One Satellite is the Test Mass: e.g. CHAMP, GRACE or GOCE**

Illustration ©ESA

# Observing Gravity Field from Space

## Satellite-to-Satellite Tracking – Low Low Case

Observation of the medium Frequencies by Analysis of Orbit Perturbations

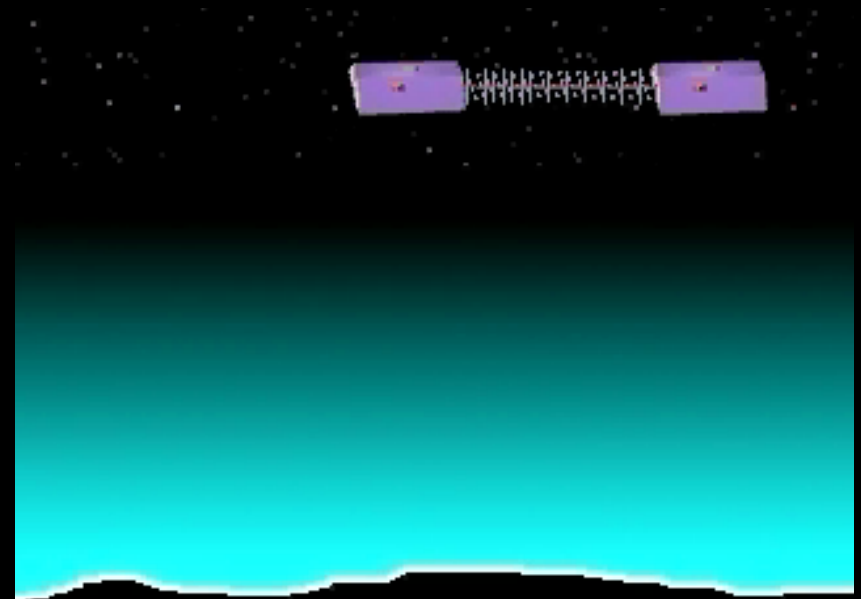
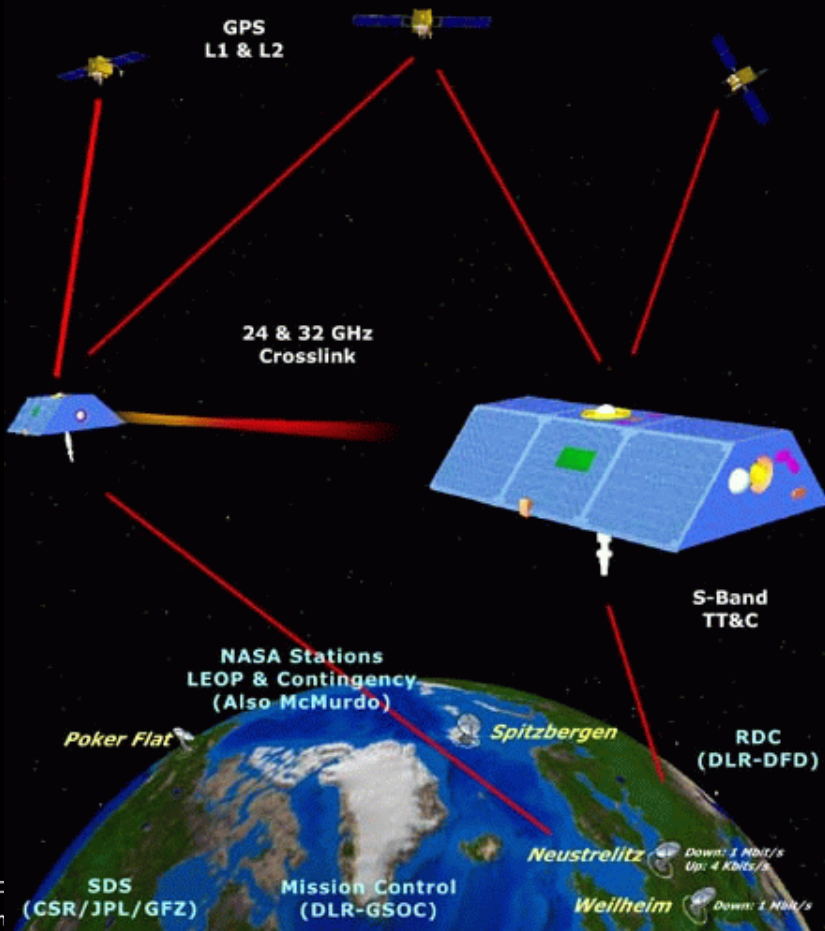


Illustration & Animation ©CSR

**Two Satellites are the Test Masses  
observing each other: e.g. GRACE**

# Observing Gravity Field from Space

## Gravity Gradiometry

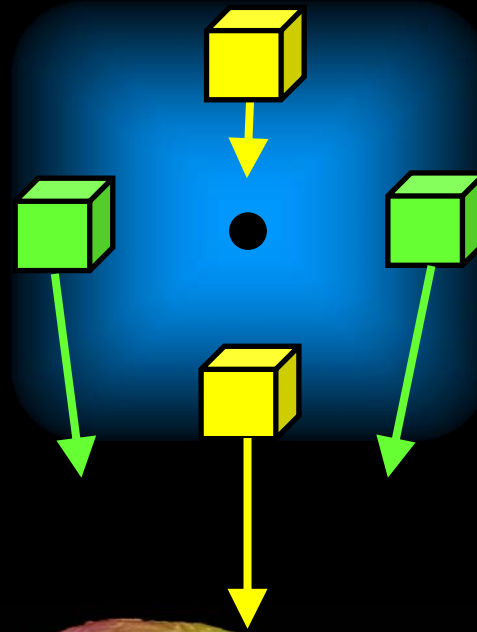
Observation of the medium  
to high Frequencies

### Observation

6 Accelerometers  
measure Accelerations in  
3 Directions.

Measurement Accuracy:  
 $10^{-12} \text{ ms}^{-2}$

Common-Mode Accelerations  
By Computation of Mean  
Value of Accelerations  
along 1 Gradiometer Arm.



### Differential Mode Accelerations

By Subtraction of  
Accelerations along 1  
Gradiometer Arm.

### Gravity Gradients

Divide differential  
Accelerations by Arm  
Length and correct for  
rotational Accelerations

**The Satellite directly observes 2nd Derivatives of the Gravity Potential: e.g. GOCE**

# Computational Challenges

## Observations & Parameters

### Number of Observations for Satellite Instrumentation:

Space GPS Receiver: 0.1 - 1 Hz, 12 channels	Accelerometer: 0.2 - 1 Hz
Intersatellite Microwave Link: 0.2 Hz	Gradiometer: 1 Hz
high-low SST: 1 Mio./day	Gradiometry: 260.000/day
low-low SST: 30.000/day	Surface Data: up to 100 Mio points

### Number of Parameters depends of chosen maximum degree of SHS:

$$n_{Par} = (N + 1)^2$$

For high-low SST – CHAMP:	$N = 80$	$n_{Par} = 6.561$
For low-low SST – GRACE:	$N = 180$	$n_{Par} = 32.761$
For Gradiometry – GOCE:	$N = 250$	$n_{Par} = 63.001$
For Combination – Sat. & Surface Data:	$N = 2160$	$n_{Par} = 4.669.921$

# Computational Challenges

## Integration of Equation of Motion

(needed for Orbit Perturbations Analysis)

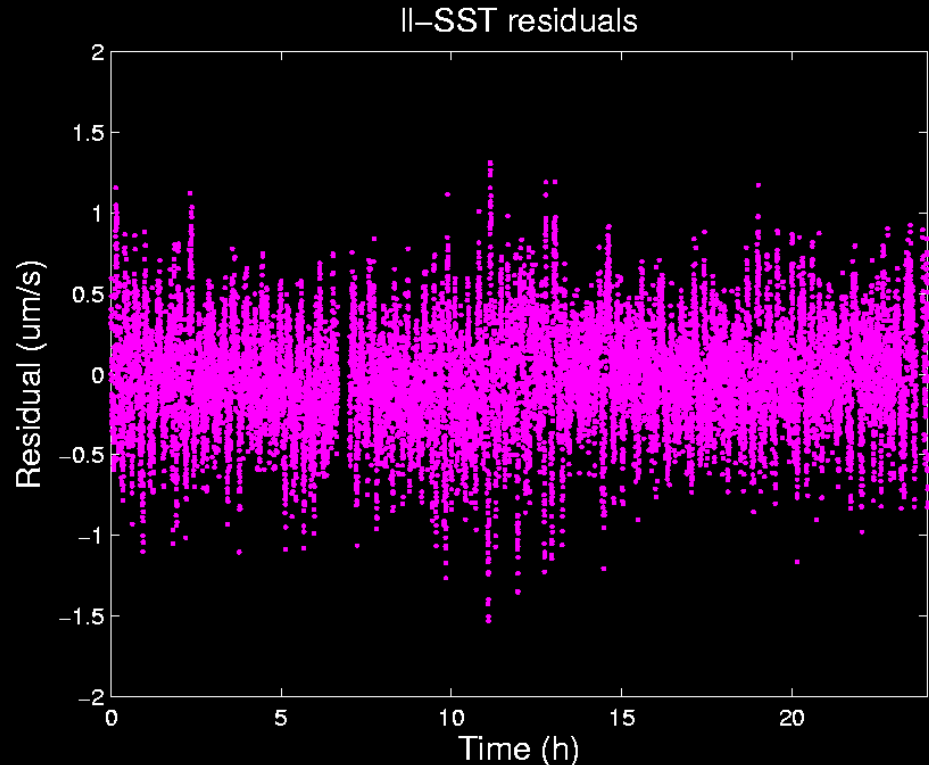
$$\ddot{\mathbf{r}} = -\frac{GM}{R^3}\mathbf{r} + \mathbf{F}_S$$

with:

- $\mathbf{r}$  Position vector in inertial frame
- $\ddot{\mathbf{r}}$  Acceleration vector in inertial frame
- $GM$  Gravity constant times Earth mass
- $R$  Distance from geocenter
- $\mathbf{F}_S$  Disturbing force vector
  - anomalous Earth gravity field
  - direct and indirect tidal forces
  - Non-gravitational (surface) forces:  
atm. drag, solar radiation, Albedo

Numerical Integrator Requirements:

- for position observations [mm]: 11 digits
- for inter-satellite range obs. [0.1 $\mu$ m]: 15 digits



# Computational Challenges

## Least Squares - Accumulation of Normal Equations

Least Squares Adjustment:

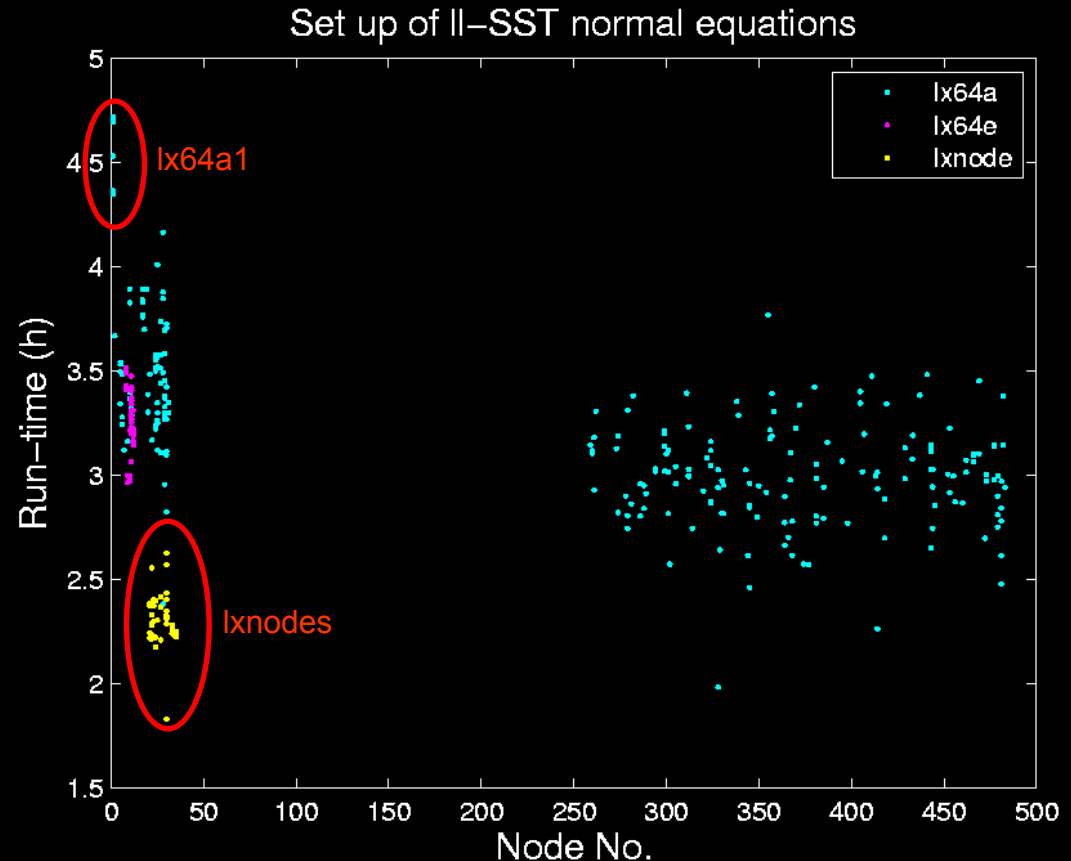
$$\mathbf{Ax} = \mathbf{l} + \mathbf{v}$$

$$E\{\mathbf{v}\} = 0 \quad \mathbf{v}^T \mathbf{Pv} = \min.$$

$$\mathbf{x} = \left( \mathbf{A}^T \mathbf{PA} \right)^{-1} \mathbf{A}^T \mathbf{Pl}$$

$$\mathbf{Q}_{xx} = \sigma^2 \left( \mathbf{A}^T \mathbf{PA} \right)^{-1}$$

$$\sigma^2 = \frac{\left( \mathbf{v}^T \mathbf{Pv} \right)}{n_{Obs.} - n_{Par}}$$



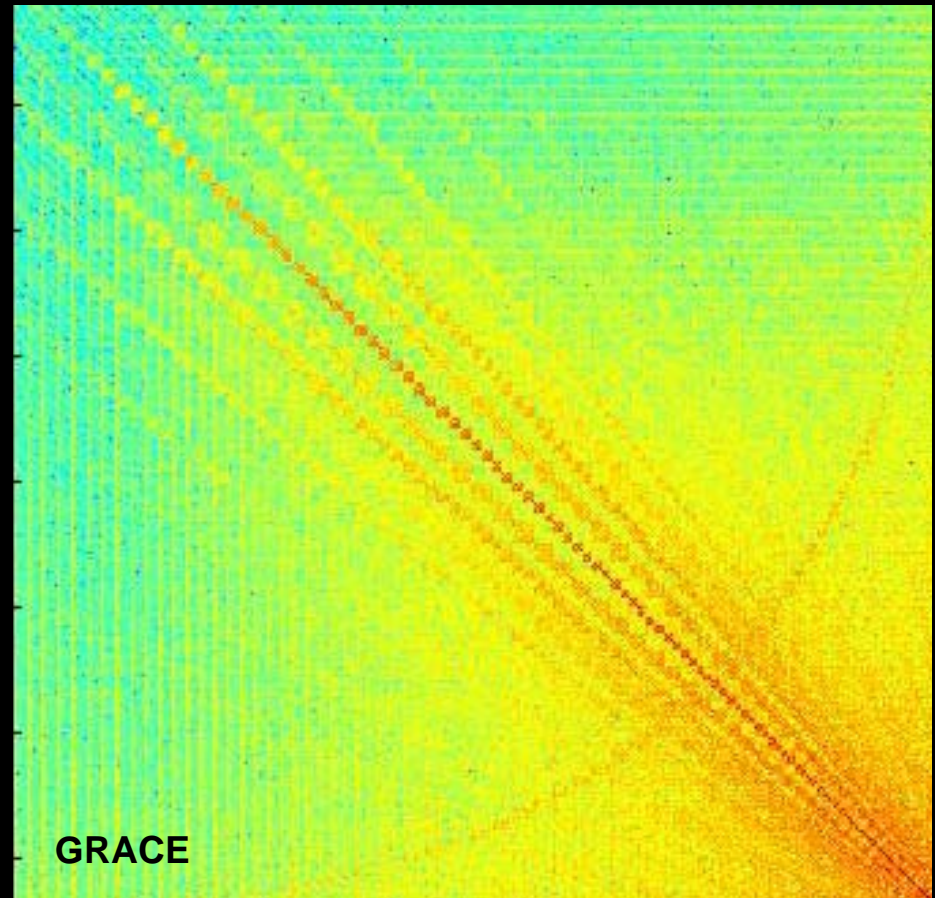
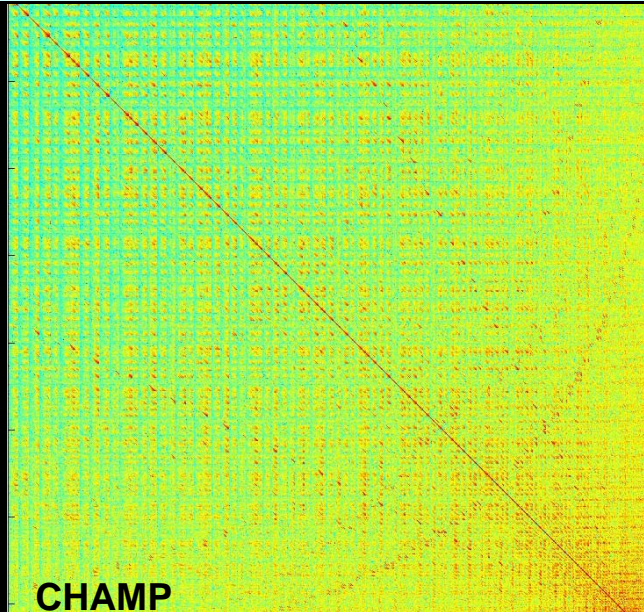
# Computational Challenges

## Least Squares – Solution/Structure of Normal Equations

Size of System to be inverted:

<b>N</b>	No. Matrix Elements	Matrix Size (8Byte)
80	$43 \times 10^6$	350 MB
180	$1 \times 10^9$	8.5 GB
250	$4 \times 10^9$	32 GB
2160	$22 \times 10^{12}$	175 TB

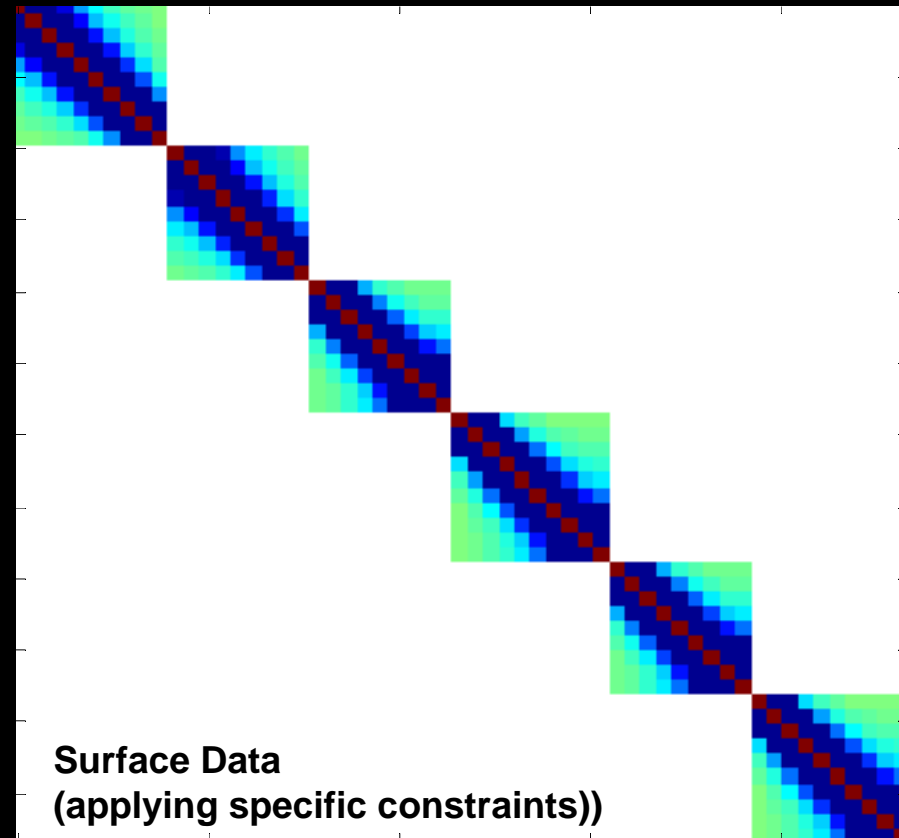
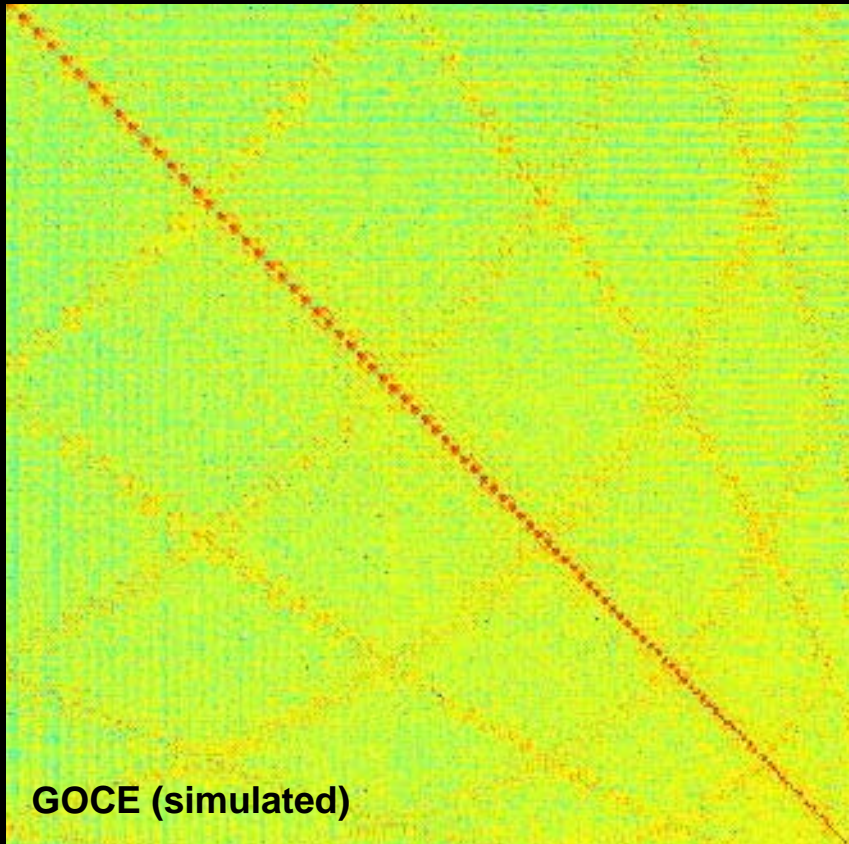
$$Q_{xx} = \sigma^2 (A^T P A)^{-1}$$



# Computational Challenges

## Least Squares – Solution/Structure of Normal Equations

$$Q_{xx} = \sigma^2 (A^T P A)^{-1}$$



# Summary & Conclusions

- **Satellites observing of the Earth Gravity Field from Space are providing a huge amount of data.**
- **The number of parameters to be determined strongly depends on the resolution (sensitivity) of the field to be determined.**
- **The global Earth gravity field is determined by an inverse process, which requires the setup and solution of large equation systems.**
- **For highest resolutions using surface observations simplifications have to be introduced in order to reduce the matrix size.**
- **The LINUX Cluster heavily is used for CHAMP and GRACE processing. It is the workhorse for our gravity field analysis.**
- **For GOCE we envisage additional computational demands. For this reason LRZ will host a „LINUX Geodesy Cluster“ currently under procurement.**